

# Kepler Problem

(1)

$$S = -Et + p_{\phi} \phi + \int dr \sqrt{2m \left( E - \frac{\alpha}{r} \right) - \frac{p_{\phi}^2}{r^2}}$$

a)  $\frac{\partial S}{\partial p_{\phi}} = \text{const} \equiv \phi_0 = 0$ , without loss of generality

$$0 = \phi - \int \frac{p_{\phi} \frac{dr}{r^2}}{\sqrt{2m \left( E - \frac{\alpha}{r} \right) - \frac{p_{\phi}^2}{r^2}}}$$

$$\phi = \cos^{-1} \frac{\frac{p_{\phi}}{r} - \frac{m\alpha}{p_{\phi}}}{\sqrt{2mE + \frac{m^2\alpha^2}{p_{\phi}^2}}} \quad \text{or} \quad \frac{p_{\phi}}{r} = 1 + \varepsilon \cos \phi$$

$$p = \frac{p_{\phi}^2}{m\alpha}, \quad \varepsilon = \sqrt{1 + \frac{2E p_{\phi}^2}{m\alpha^2}}$$

b)  $\frac{\partial S}{\partial E} = \text{const} = -t + \int \frac{m}{\sqrt{2m \left( E - \frac{\alpha}{r} \right) - \frac{p_{\phi}^2}{r^2}}} dr$

(2)

# Kepler Problem

Consider  $E = -\frac{md^2}{2p\phi^2}$ ,  $\mathcal{E} = 0$ ,  $r = p$   
circular motion

$$S = \frac{md^2}{2p\phi^3} t + p\phi\phi$$

$$\frac{\partial S}{\partial p\phi} = \text{const} = -\frac{md^2}{p\phi^3} t + \phi$$

const  $\equiv \phi_0 = 0$ , without loss of generality

$$\phi = \frac{md^2}{p\phi^3} t \quad r = p = \frac{p\phi^2}{md}, \quad d^2 = \frac{p\phi^4}{m^2 r^2}$$

$$= \frac{p\phi}{mr^2} t$$

Alternatively

$$S = -Et + \sqrt{-\frac{md^2}{2E}} \phi$$

$$\frac{\partial S}{\partial E} = \text{const} = 0 = -t + \sqrt{\frac{md^2}{2}} \frac{1}{2(-E)^{3/2}} \phi$$

$$= -t + \frac{p\phi^3}{m d^2} \phi, \text{ as before}$$