

Welcome to Class 4:
Our Solar System
(and a bit of cosmology)

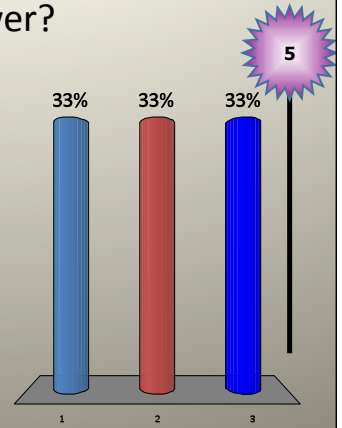


What is the difference between:
dark ENERGY and dark MATTER?

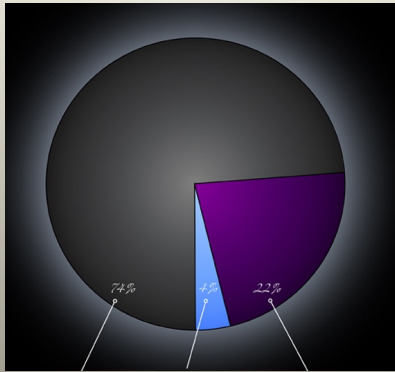
Is Earth unique, or might Earth-like objects
form elsewhere?

Which learning objectives do you find
most difficult to answer?

1. Compare/Contrast Dark Matter/Energy give evidence of them.
2. Label/Describe the four stages of Nebular Theory
3. List 5 pieces of evidence for Nebular Theory from observed solar system properties.

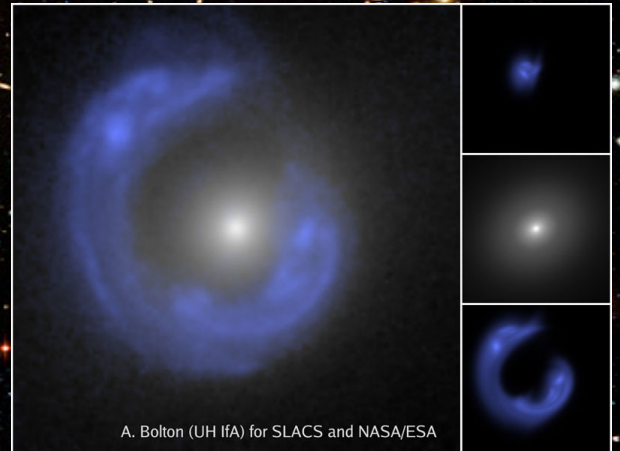


What is our Universe made up of?



Dark Energy Normal Matter Dark Matter

What evidence is there for dark matter?

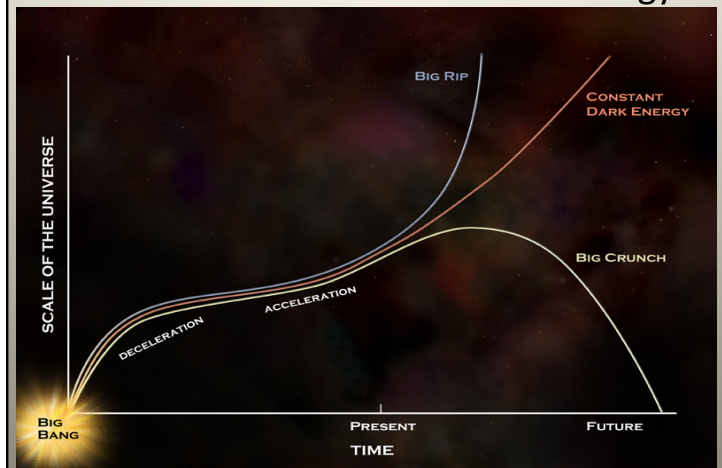


A. Bolton (UH IfA) for SLACS and NASA/ESA

Dark matter (blue) vs.
'dark' Normal Matter (pink)

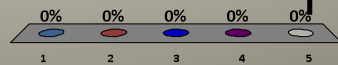


What evidence is there for Dark Energy?



Discuss then answer with your PRS:
Which is not a 'Terrestrial' planet?

1. Saturn
2. Mercury
3. Mars
4. Venus
5. Earth



Working in groups, fill in this chart in your notes:

Characteristics	Terrestrial Planets	Jovian Planets
Location in Solar System		
Composition		
Mass (high or low?)		
Radius (large or small?)		
Distances between neighboring planets		

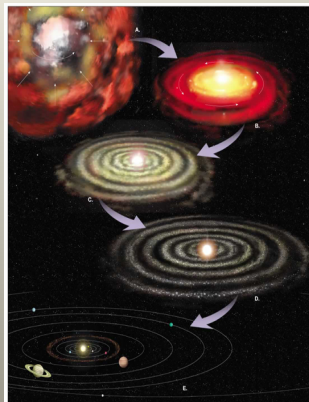
The 'Nebular Theory' of planet formation

Used to explain all observations:
the *differences* & the *similarities*
between planets.

Basically:

Our solar system was born
from the gravitational collapse
of an interstellar cloud, or
nebular, of gas and dust.

This cloud/nebula is referred to
as the 'solar nebula'.



A massive cloud collapses under its own gravity
and spins up due to conservation of angular
momentum. This creates a flat, spinning disk,
with most of the mass at the center.

1. Contraction

Contraction: As it contracts, the cloud heats,
flattens, and spins faster, becoming a spinning
disk of dust and gas.

A large, diffuse interstellar gas cloud
(solar nebula) contracts due to gravity.

The Sun will be born in the center.

Planets will form in the disk.

Which characteristics does this stage explain?

The nebula center remains hot and dense,
the outer regions cool and tenuous

Condensation: Hydrogen and helium remain
gaseous, but other materials can condense into
solid "seeds" for building planets.

2. Condensation



Which characteristics does this stage explain?

Planetesimals accrete to build up planets from
the local material available

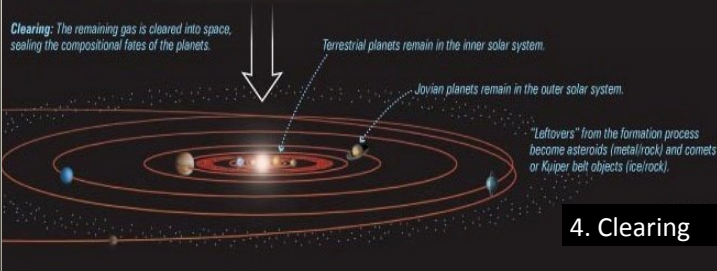
Accretion: Solid "seeds" collide and stick together.
Larger ones attract others with their gravity,
growing bigger still.

3. Accretion



Which characteristics does this stage explain?

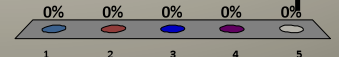
The rest of the smaller debris is blown away once the Sun 'turns on'



Which characteristics does this stage explain?

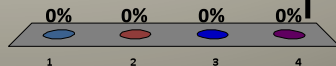
PRS: What Solar System characteristic is not explained by the Nebular Theory?

1. Planets orbit in the same direction
2. The inner planets are rocky
3. The sun lies at the center
4. The planets lie in a flat plane
5. Venus's backward revolution



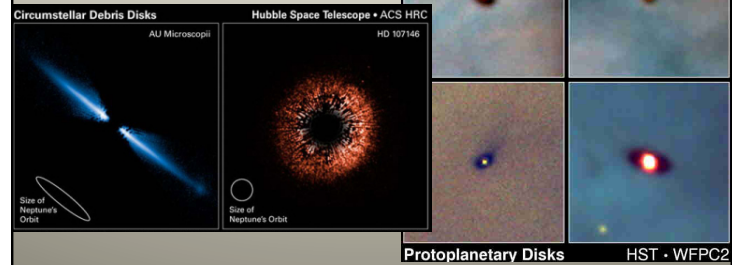
PRS: Why did the Jovian planets get so big?

1. They formed planet 'seeds' of rock, but also mixed with the very abundant ices available.
2. These massive rock/ice seeds formed larger planetesimals.
3. Their planetesimals grew very large very quickly, to capture H and He before the Sun turned on and blew it away.
4. All of the above.



What is the relevance of the 'Nebular Theory' for searching for life in the Universe?

With HST we see these disks around distant young stars.

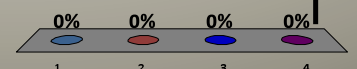


Let's test your understanding of today's learning objectives.

Please take this test alone.

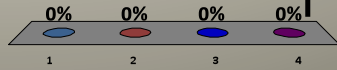
1. What is most of our Universe made of?

1. Protons & Electrons
2. Normal Matter
3. Dark Matter
4. Dark Energy



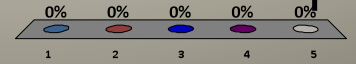
2. If the universe continues on its present course it will..

1. Expand forever
2. Collapse to a small crunch
3. Expand for a time, but then fall back to a small crunch
4. The present course is uncertain.



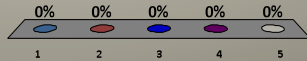
3. Which is not a stage in the Nebular Theory?

1. Contraction
2. Division
3. Condensation
4. Accretion
5. Clearing



4. In what stage do we see the two differing planetary compositions begin to form?

1. Contraction
2. Division
3. Condensation
4. Accretion
5. Clearing



5. Terrestrial and Jovian planets are similar in what characteristic?

1. Chemical Composition
2. Mass
3. Radius
4. Orbital properties

